

Asteroseismology and exoplanets: a lesson from the Sun *

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- Can we learn on exo-planets from asteroseismology?
- Metallicity of stars with planets
- The origin of metals in the Sun photosphere and helioseismology.

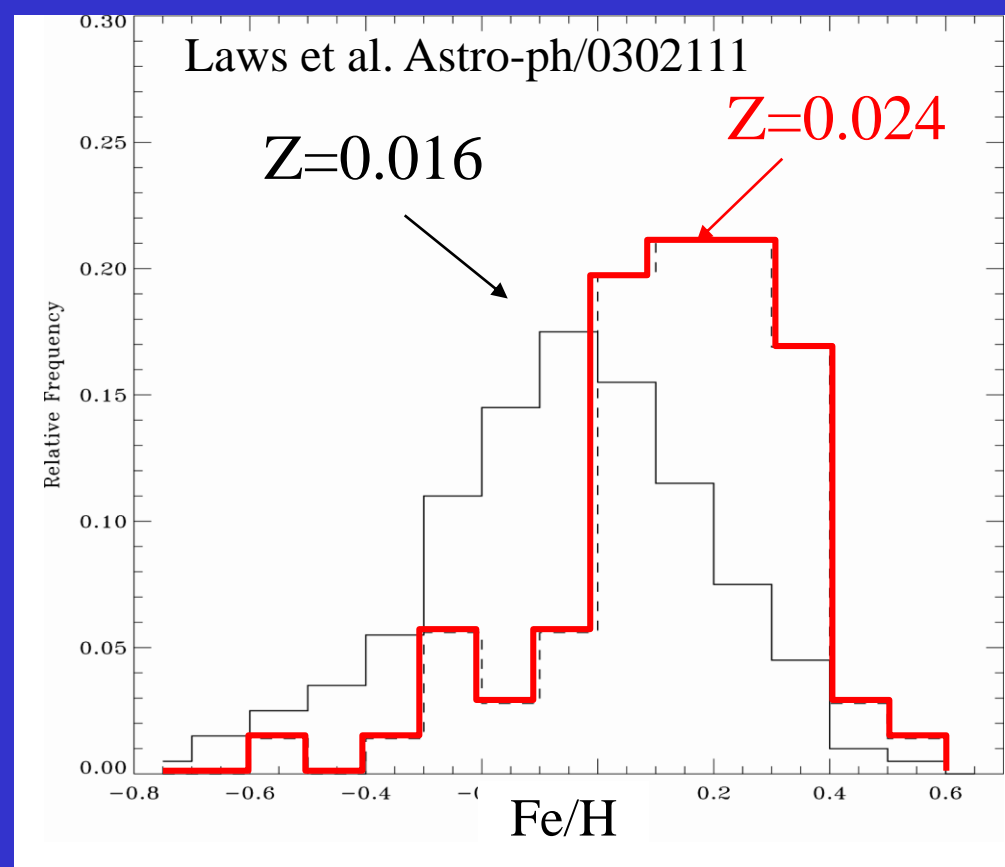
*based on: A. Zanzi, Master Thesis, Ferrara University (2003)
and discussions with Castellani Ortolani Paterno'

The Chicken and the Egg dilemma

- Stars hosting planets look more metal rich.
- The metal content in the convective envelope of a solar type star with $Z=0.016$ is

$$\Delta M_Z = 120 M_{\text{earth}}$$

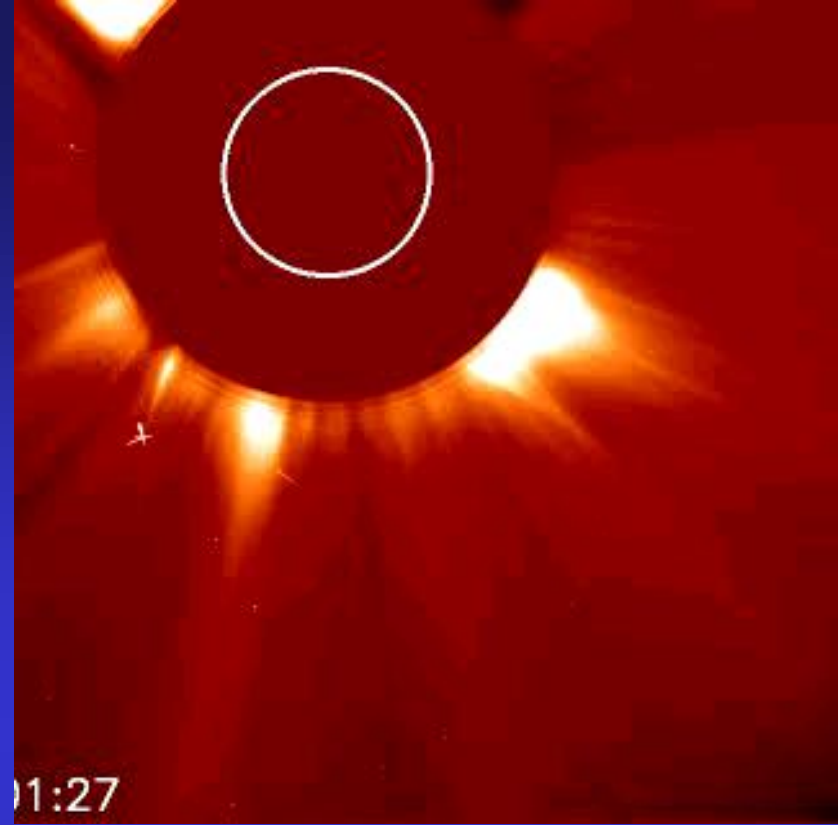
- To reach $Z=0.024$ an extra $60 M_{\text{earth}}$ are needed.
- Is the excess metallicity primordial or a result of accretion?



- Is it a surface/ bulk effect ?
 $Z_{\text{int}} < \text{or} > Z_{\text{ph}}$
- Need a look at the stellar interior: asteroseismology is the natural tool

Dirty solar models

- The Sun is a natural laboratory
- Matter is falling onto the Sun (and presumably more was falling in the past)
- The SSM (without accretion) has been accurately tested with helioseismology (& neutrinos)
- Build solar models starting with a lower metallicity and assume metal accretion occurs (shortly) after ZAMS.



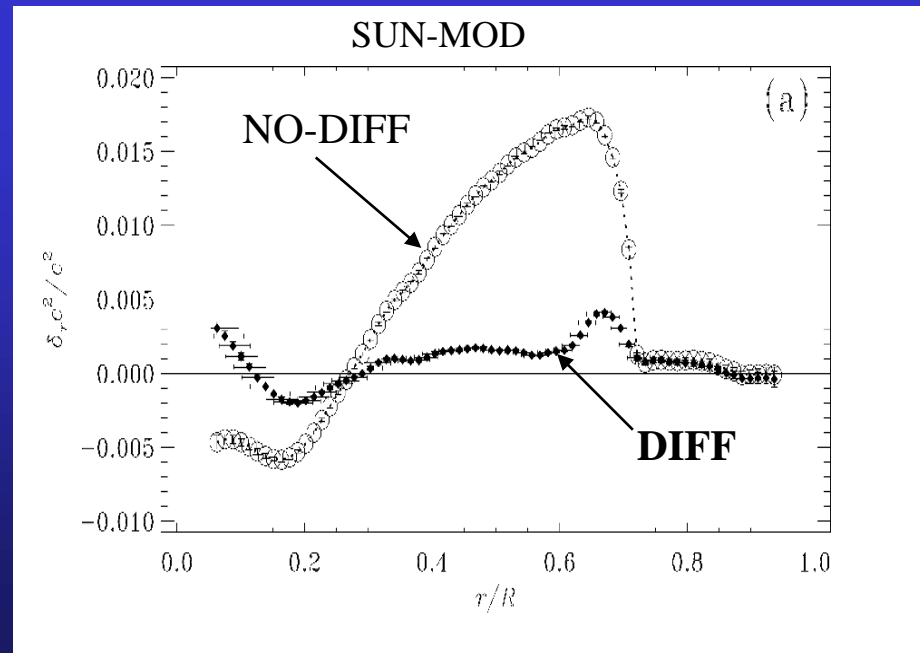
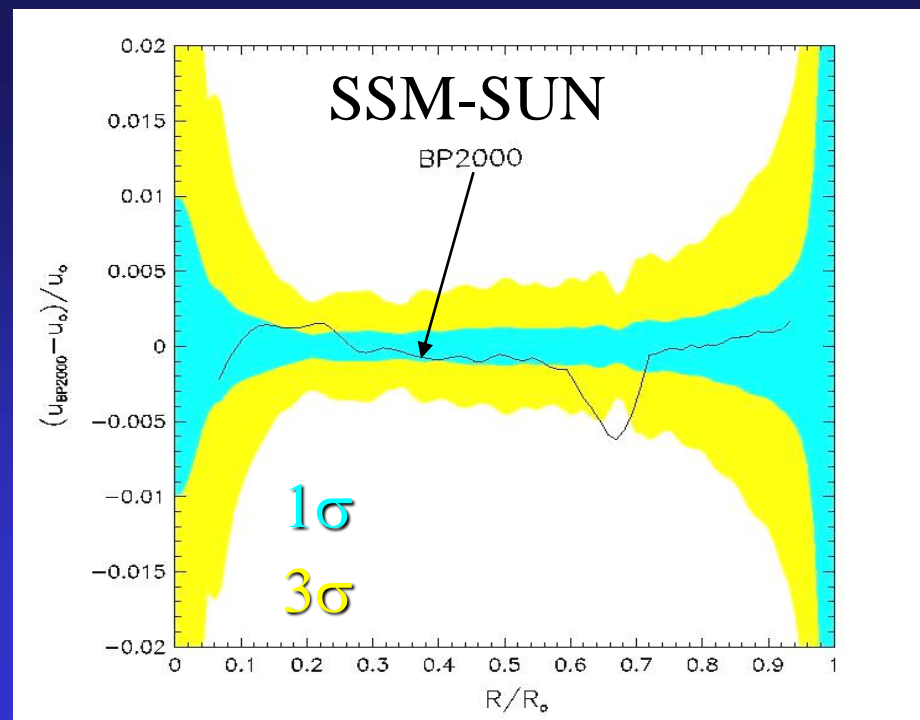
- How much material can have accreted without spoiling the agreement with helioseismology (& neutrinos)?

The lesson of diffusion

- SSM are in excellent agreement with helioseismology (for u , Y_{ph} , $R_b...$) when diffusion is included.
- The effect of diffusion is that

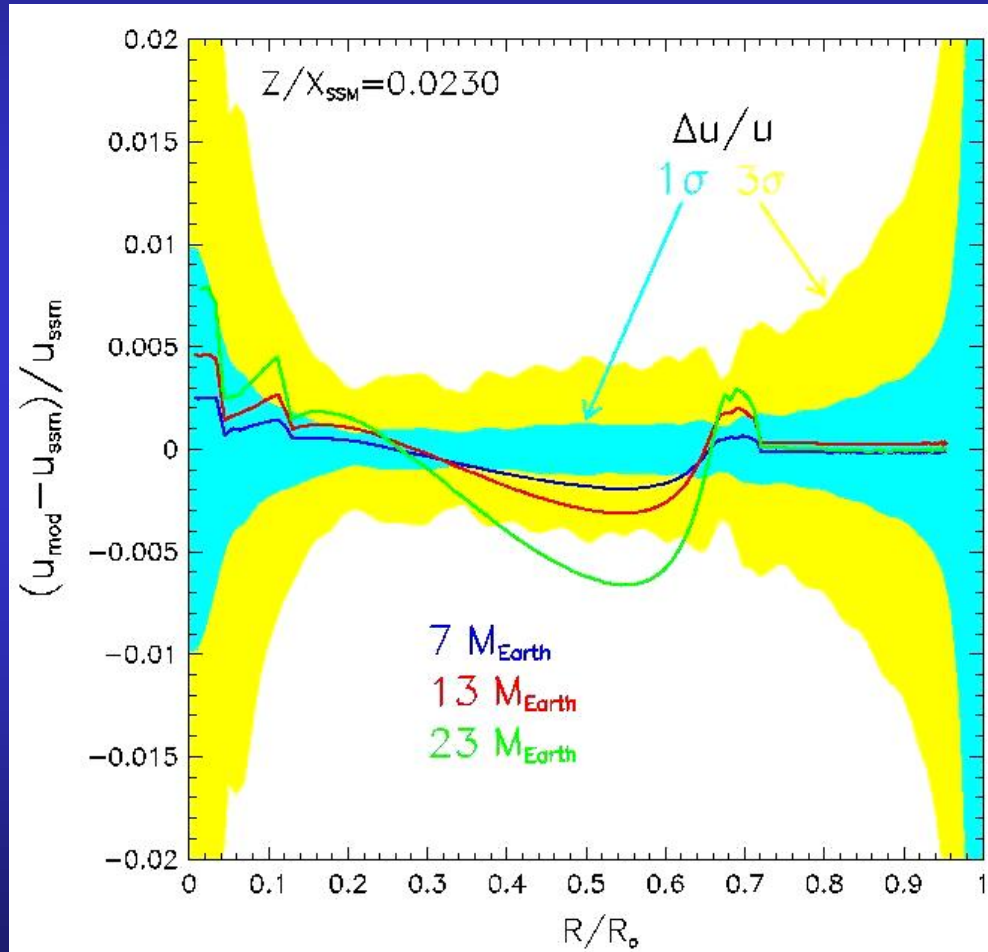
$$Z_{ph} = 0.9 Z_{int}.$$

- This means that helioseismology is sensitive to changing by 10% the metal mass in the convective envelope, i.e. to $12M_{earth}$



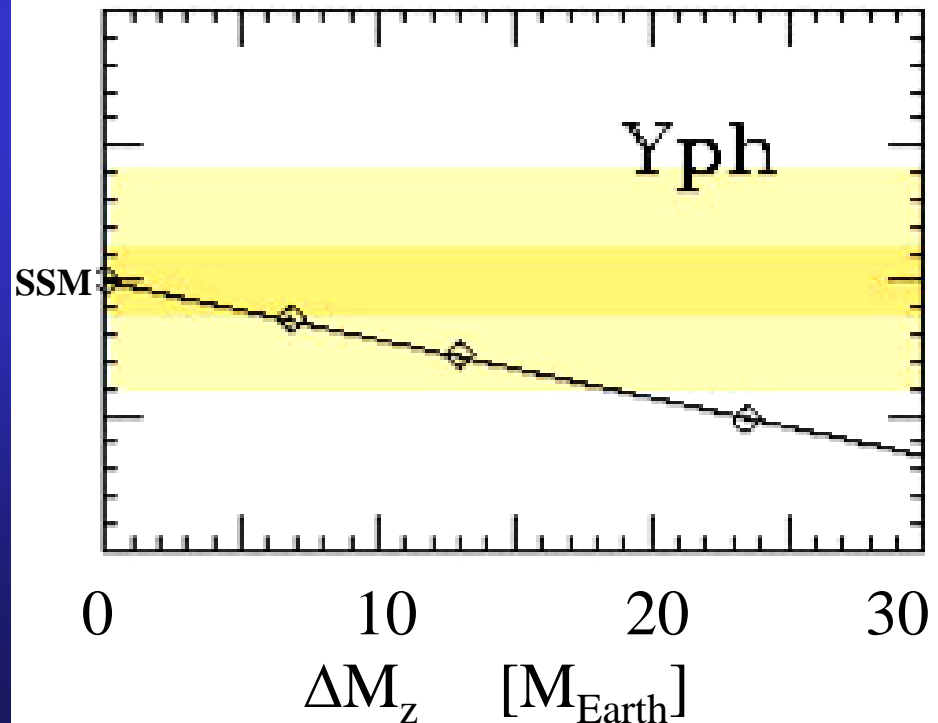
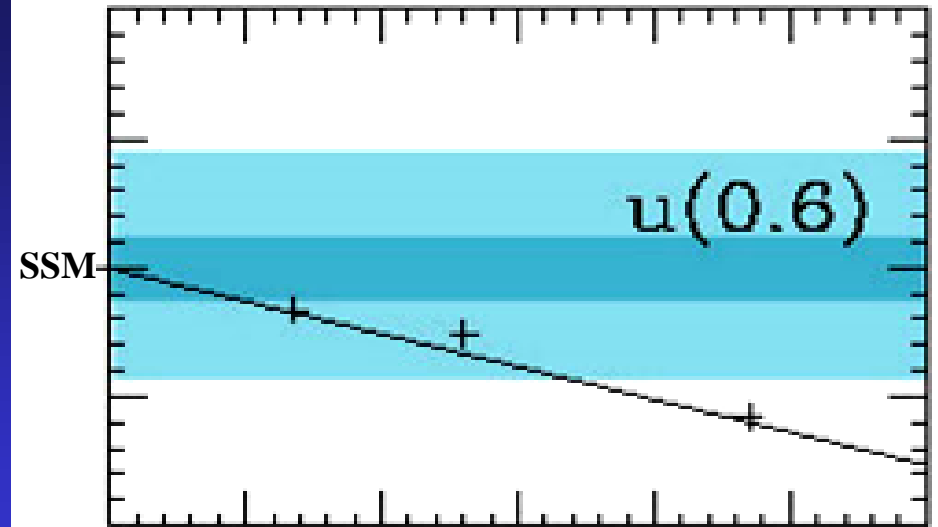
Dirty solar models

- With respect to SSM, a dirty Sun is metal poor in the interior.
- molecular weight and opacity are changed
- Temperature, sound speed, Y_{ph} (& Φ_{ν}) can be changed



Constraints from helioseismology and neutrino

- Can study the change of several observables (sound speed, Y_{ph} , R_b , T_c)
- Helioseismic observables imply:
 $\Delta M_Z < 20 M_{Earth}$ (at 3σ)
- [Boron neutrino flux gives $\Delta M_Z < 30 M_{Earth}$]



Helioseismology and Asteroseismology

- Helioseismology is sensitive to $O(10) M_{\text{earth}}$
- For asteroseismology we need sensitivity to $O(100) M_{\text{earth}}$
- We are planning simulations (which frequencies? Which accuracy ?)
- Asteroseismology can prove accretion theory:
$$Z_{\text{int}} < Z_{\text{ph}}$$
- However it cannot disprove it (if accretion occurred before ZAMS).

